

TAMA300 interferometer

● **Laser interferometer GW detector**

- Arm length: 300m
- Location: National Astronomical Observatory of Japan (Mitaka, Tokyo)

● **Purposes**

- Development of the detector capable to catch GW events in nearby galaxies
- Establishment of interferometer technologies for LCGT



Progress of TAMA300

1995-1997 Facility/Vacuum system construction

Recombined Interferometer

1999-2001 6 times of observation runs
(Total 1370 hours)

Recycled Interferometer

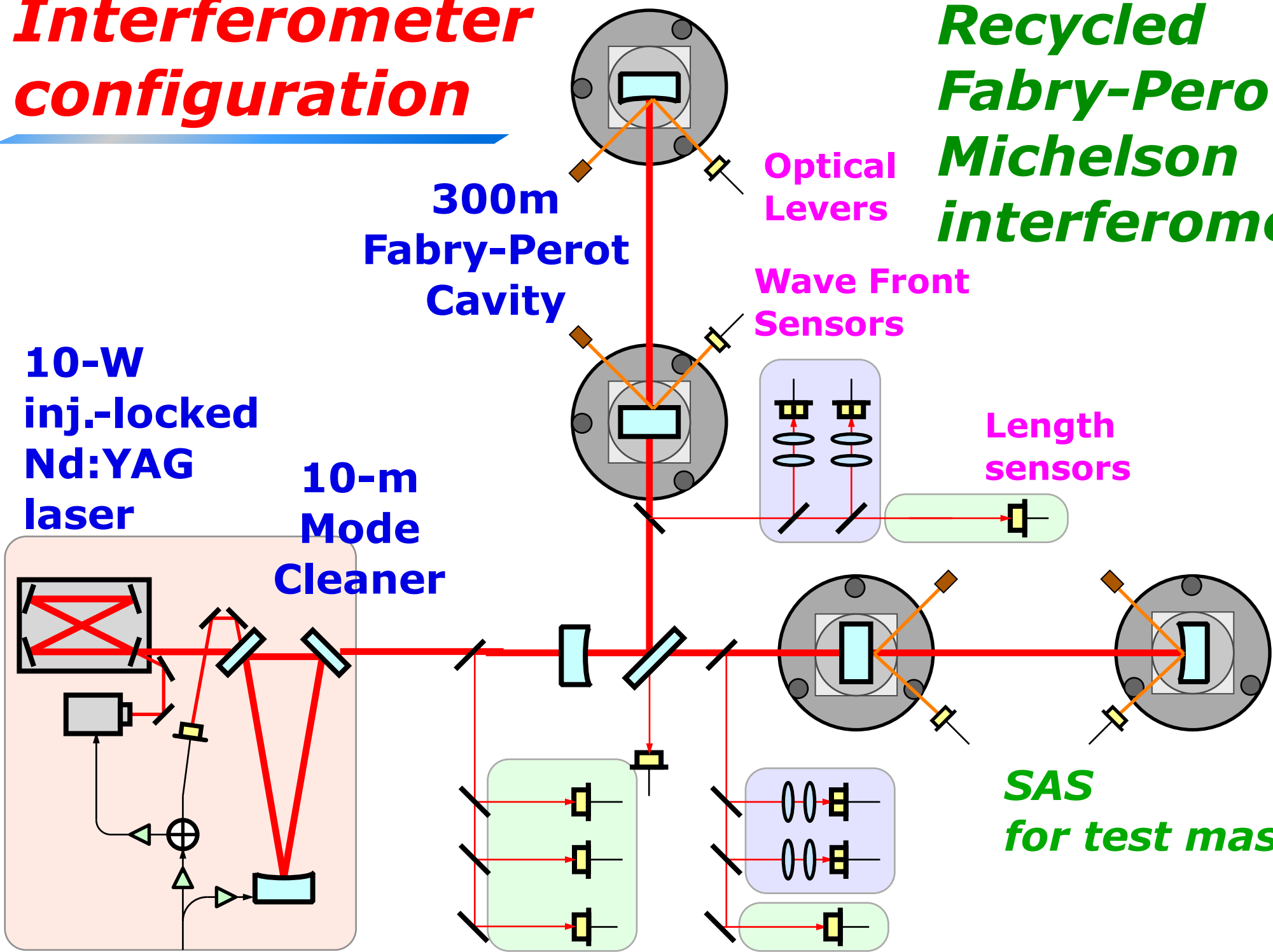
2001 Implement of power recycling
2003-2004 3 times of observation runs
(Total 1740 hours)

Seismic Attenuation System (SAS)

2005 Start installation of SAS
2007 Full interferometer lock with SAS
2008 Sensitivity improvement

Interferometer configuration

Recycled Fabry-Perot Michelson interferometer



**10-W
inj.-locked
Nd:YAG
laser**

**10-m
Mode
Cleaner**

**300m
Fabry-Perot
Cavity**

**Optical
Levers**

**Wave Front
Sensors**

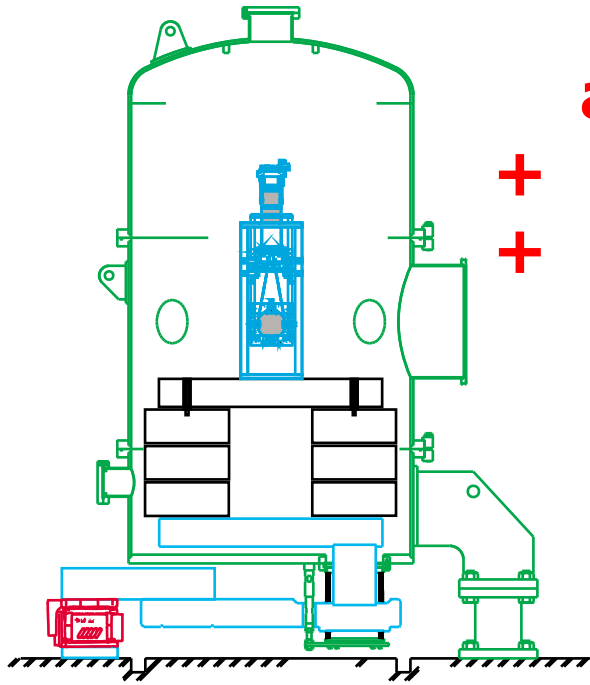
**Length
sensors**

**SAS
for test masses**

Current focus

● Establishment of detector operation with SAS

Replacement of the vibration isolation system



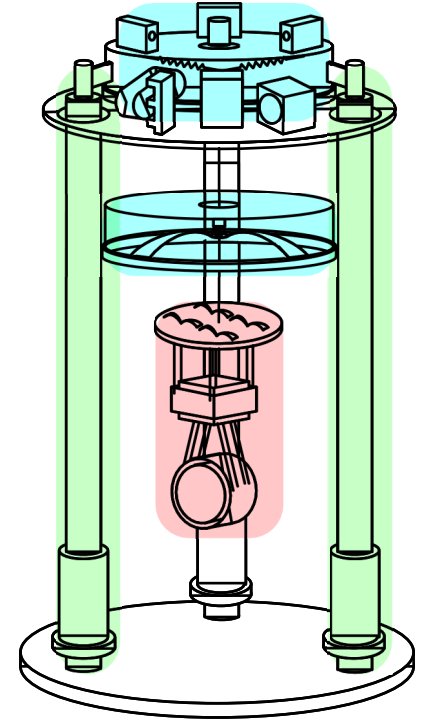
**pneumatic
active isolator**

+ stack

+ double pendulum



**inverted pendulum
+ vertical filter
+ multiple pendulum**



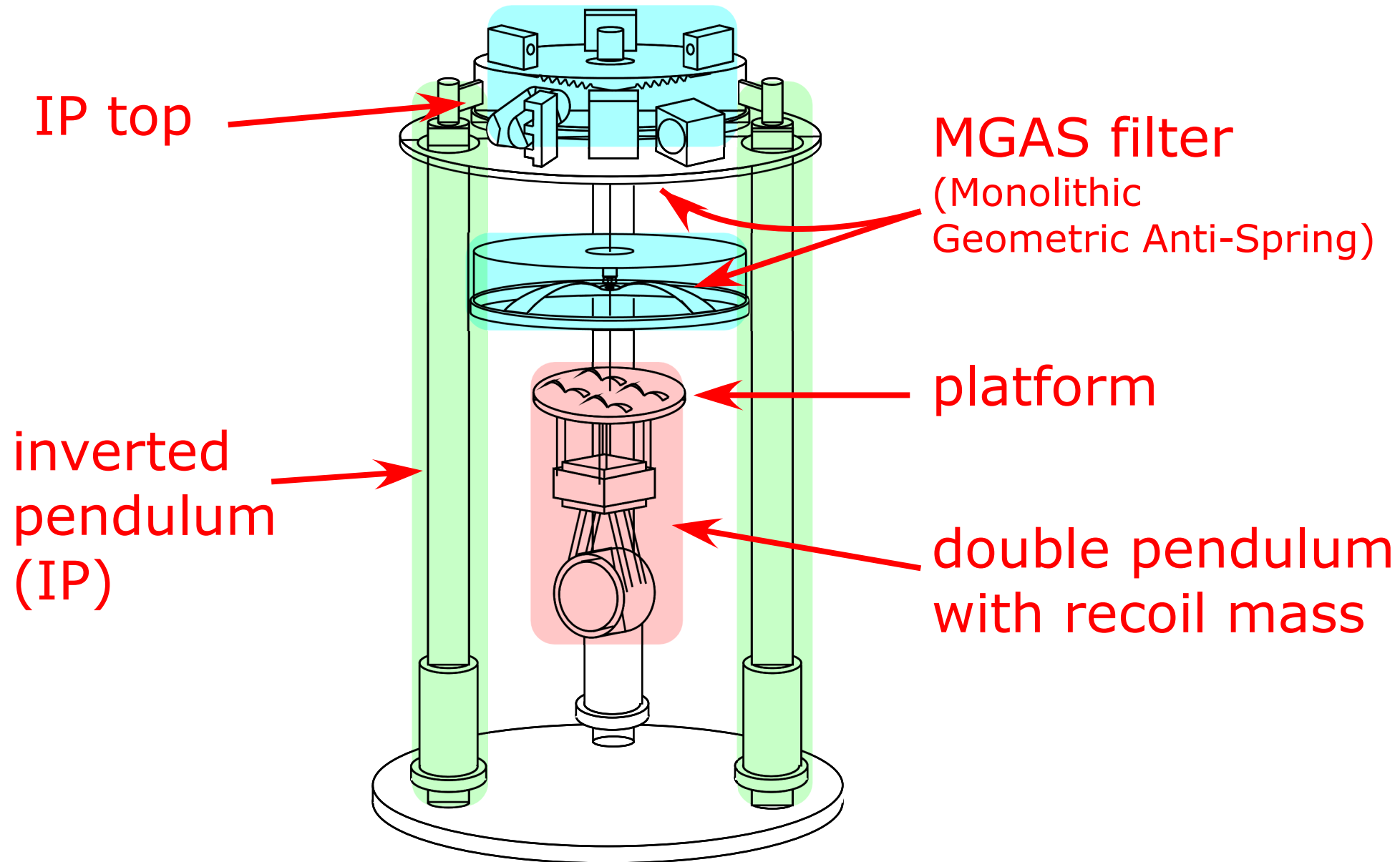
To realize the improvement

- Optimization of SAS control
- Optimization of interferometer control
- Application of digital control system
 - => Enables the complex servo system
 - => High level automatization of the operation

Seismic Attenuation System

- **Structure of SAS**

Multiple pendulum suspended from IP



Seismic Attenuation System

● Vibration Isolation

Passive isolation with soft springs + active damping

Torsional

Tortion Pendulum

$f \sim 40\text{mHz}$

Inverted Pendulum

$f \sim 500\text{mHz}$

Horizontal

Inverted Pendulum

$f \sim 30\text{mHz}$

Multiple Pendulum

$f \sim 650\text{mHz}$

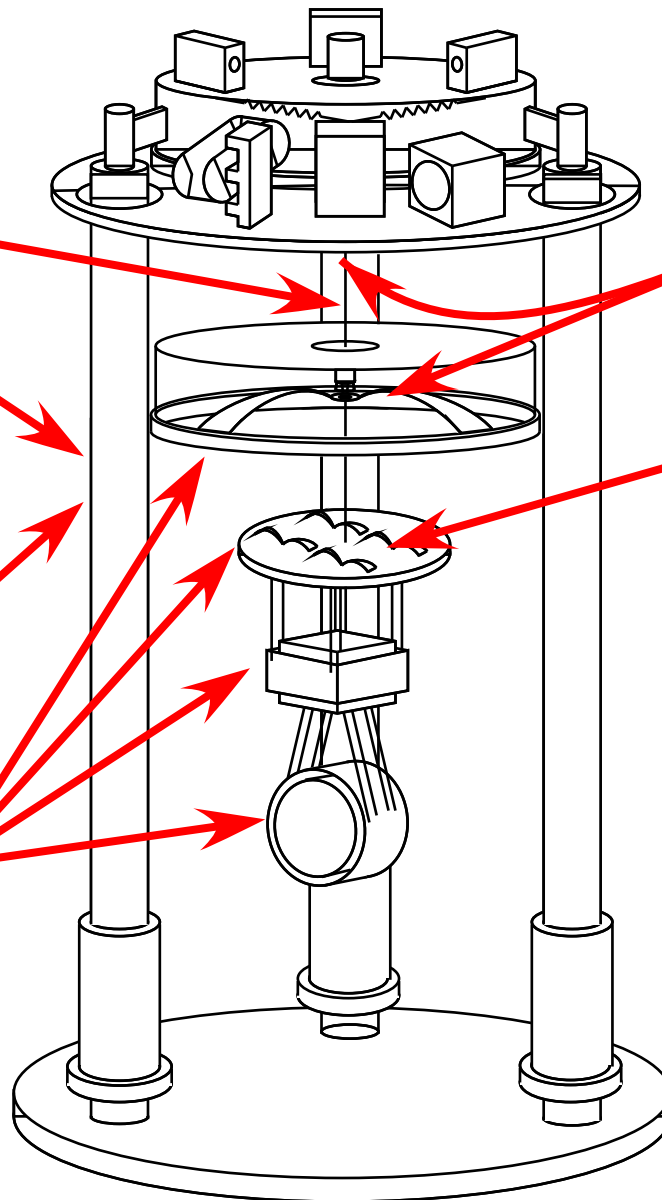
Vertical

MGAS Filter

$f \sim 50\text{mHz}$

MiniGAS Filter

$f \sim 1.5\text{Hz}$



Seismic Attenuation System

● Active Control of SAS

Local control stabilize the mirror motion

=> to enable lock of the interferometer

Local control

IP Position

Sensor: LVDT

Bandwidth: $\sim 60\text{mHz}$

IP Inertial damping

Sensor: Accelerometer

Bandwidth: $60\text{m}\sim 2\text{Hz}$

Tortion damping

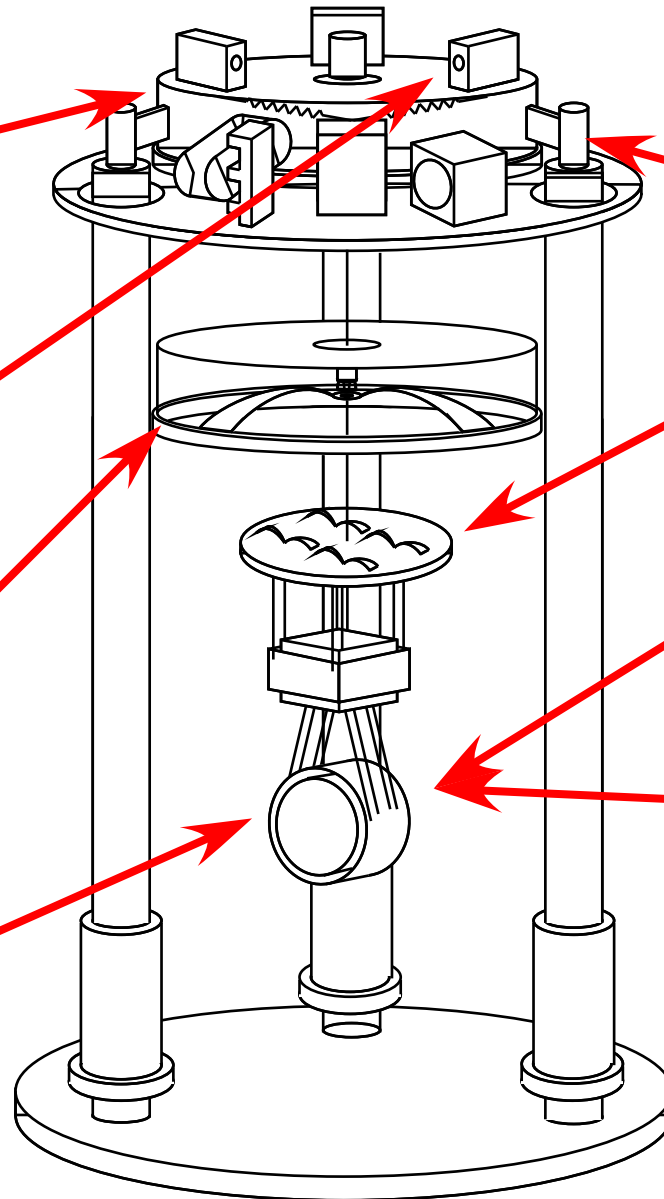
Sensor: Photo Sensor

Bandwidth: 40mHz

Test mass servo

Sensor: Optical Lever

Bandwidth: $\sim 2\text{Hz}$



Global control

IP Position

Bandwidth: $\sim 10\text{mHz}$

Plat form

Bandwidth: $\sim 10\text{mHz}$

Test mass (angular)

Bandwidth: $< 3\text{Hz}$

Test mass (Length)

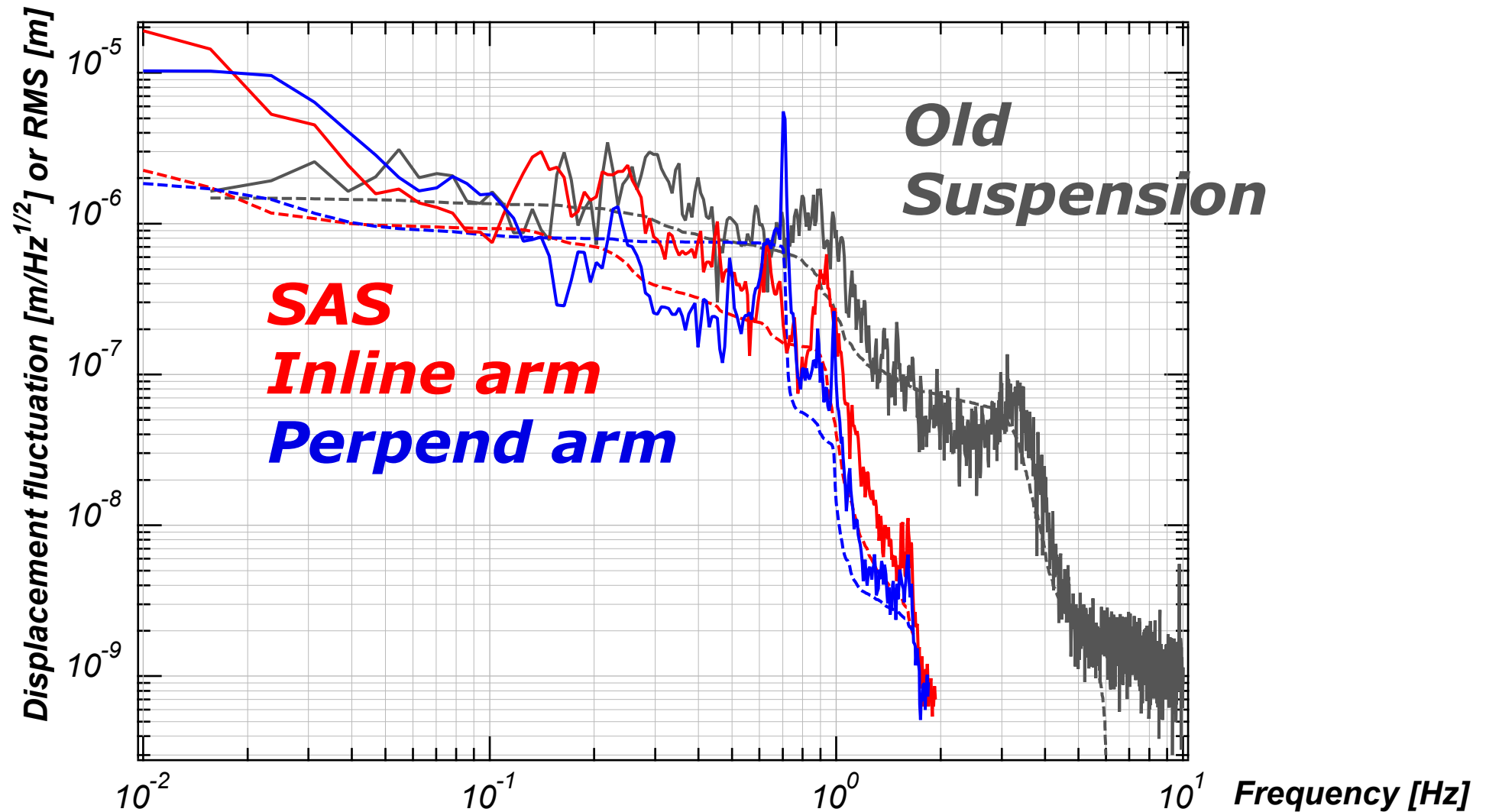
Bandwidth: $< 1\text{kHz}$

Performance of SAS

● Low frequency Legth Fluctuation of 300-m arm

Comparison with the previous suspension system

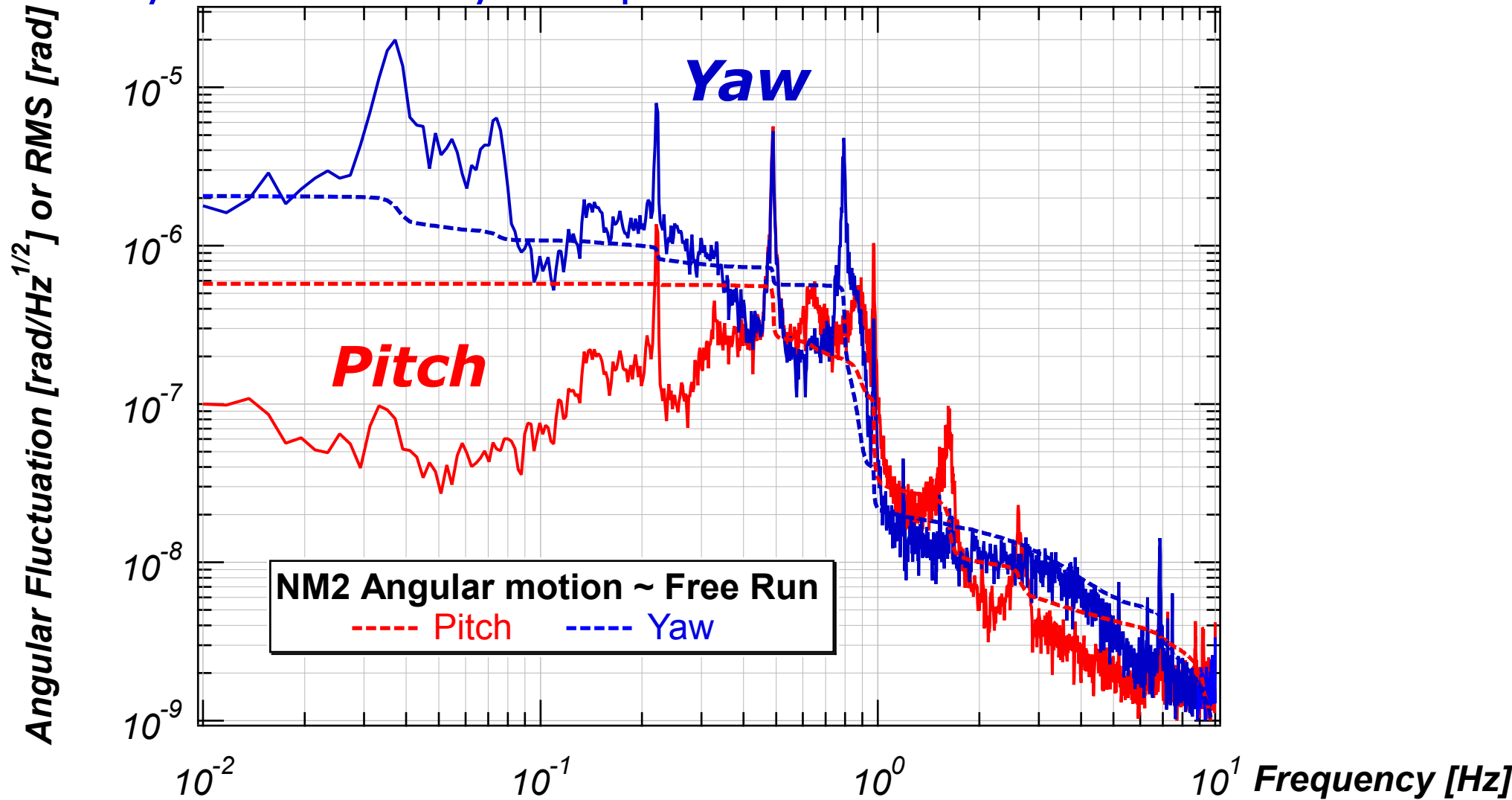
=> **improvement above 0.1Hz was confirmed**



Performance of SAS

● Test mass angular motion (Free run)

locally measured by an optical lever



Fluctuation Power concentrated on the low freq

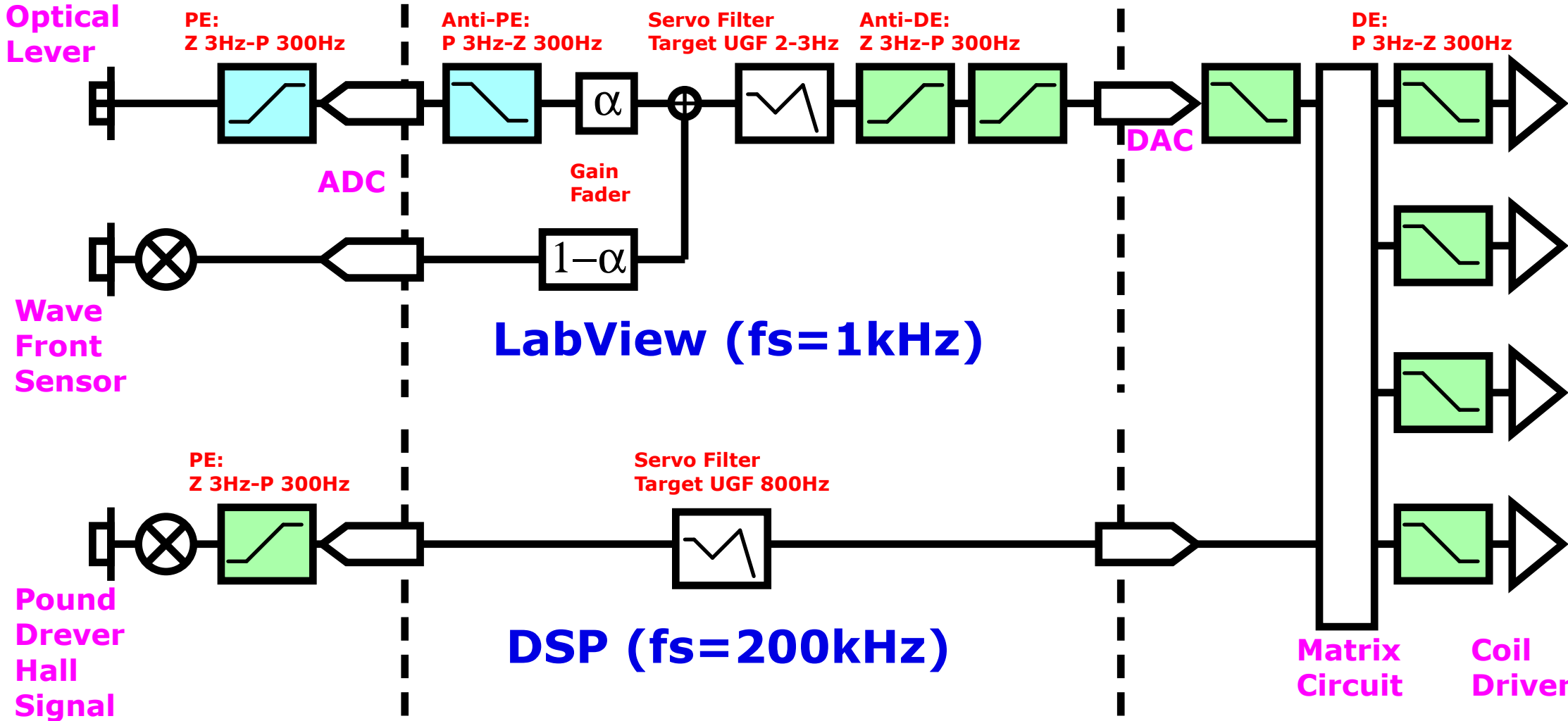
$\theta_{\text{RMS}} \sim 30 \text{ nrad} (f > 1 \text{ Hz})$

Digital Control System

Analog

Digital

Analog



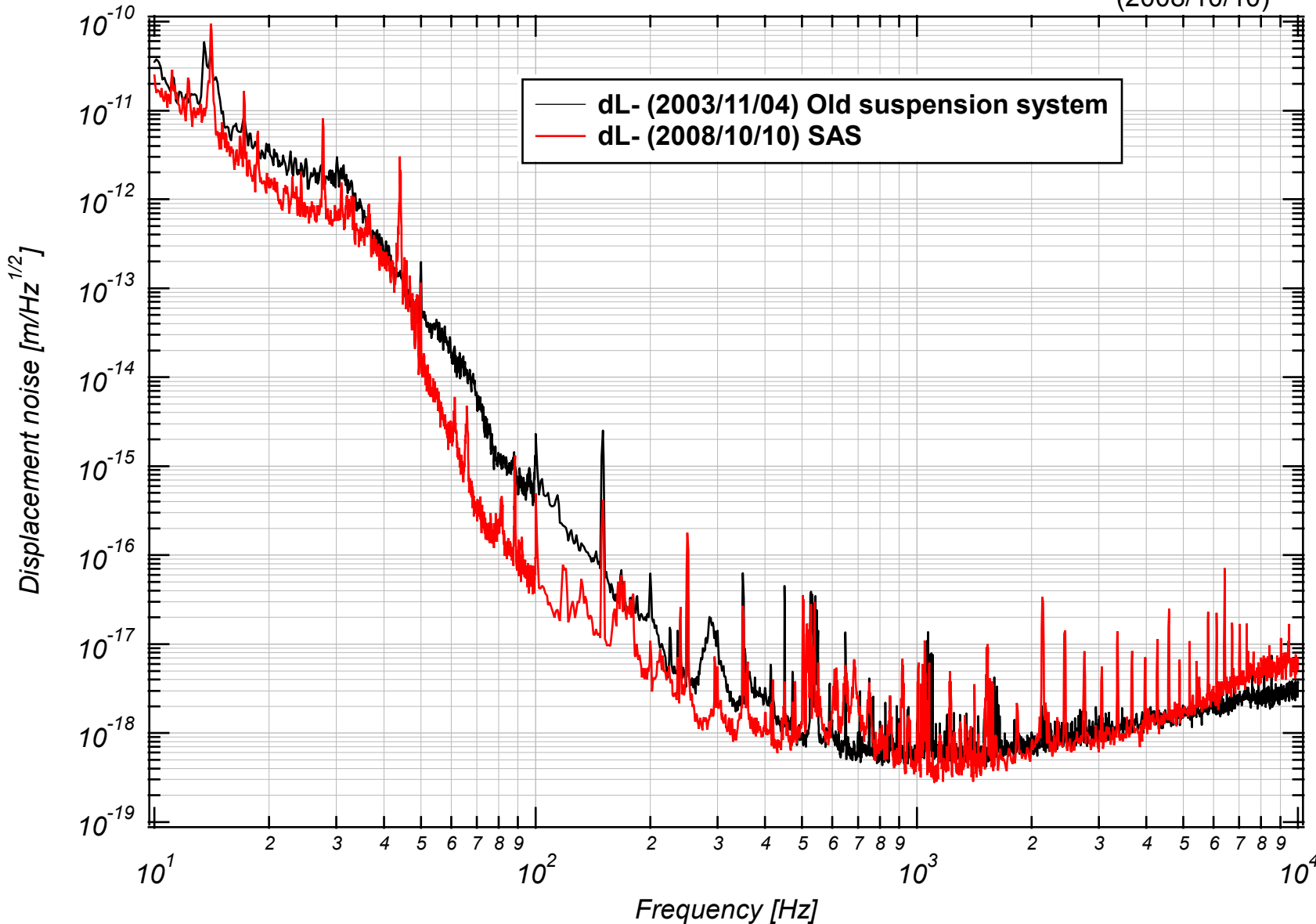
Sensitivity

● Sensitivity improvement achieved

So far, improvement below 150Hz was confirmed

TAMA300 Displacement Sensitivity

(2008/10/10)



floor level
@1kHz

displacement
 $dL=4 \times 10^{-19}$
[m/rtHz]

strain
 $h=1.3 \times 10^{-21}$
[/rtHz]

Noise budget

- Estimated contributions of the various noise sources

TAMA300 noise budget

(2008/10/10)

